



On the regularization of extremal three-point functions involving giant gravitons



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ABSTRACT

In the AdS_5/CFT_4 set-up, extremal three-point functions involving two giant 1/2 BPS gravitons and one point-like 1/2 BPS graviton, when calculated using semi-classical string theory methods, match the corresponding three-point functions obtained in the tree-level gauge theory. The string theory computation relies on a certain regularization procedure whose justification is based on the match between gauge and string theory. We revisit the regularization procedure and reformulate it in a way which allows a generalization to the ABJM set-up where three-point functions of 1/2 BPS operators are not protected and where a match between tree-level gauge theory and semi-classical string theory is hence not expected.

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After the successful application of integrability techniques to the planar spectral problem of the AdS_5/CFT_4 set-up [1], the calculation of three-point functions in the same set-up has attracted renewed attention with some recent highlights being the conjecture of an all loop formula for three-point functions of single trace operators in certain sub-sectors of $\mathcal{N} = 4$ SYM [2] and the formulation of certain integrability axioms for the cubic string theory vertex [3].

We will be considering three-point functions which do not belong to the class of three-point functions considered in the above references. Our three-point functions involve giant gravitons which in the string theory language correspond to higher dimensional D- or M-branes wrapping certain submanifolds of the string theory background and which in the gauge theory picture are represented by specific linear combinations of multi-trace operators, namely Schur polynomials. Remaining in the gauge theory picture, our three-point functions will involve two Schur polynomials and one single trace operator, all of 1/2 BPS type. Furthermore, the three operators will be chosen such that $\Delta_1 = \Delta_2 + \Delta_3$, where the Δ 's are the conformal dimensions of the operators. Such three-point functions are denoted as extremal three-point functions and are known to require special care in the comparison between gauge and string theory [4]. On the gauge theory side the three-point functions of interest can be calculated using techniques from zero-

dimensional field theories [5] (see also [6]) and on the string theory side they can be determined by generalizing a method developed for the calculation of heavy-heavy-light correlators [7–9] from string states to membranes [5].

In the case of the $AdS_5 \times S^5$ correspondence the 1/2-BPS nature of the operators involved implies that the three-point function considered is protected and thus should take the same value whether calculated in string theory or in gauge theory. As pointed out in [10] the need for special treatment of extremal correlators in string theory is relevant here and in [11] a regularization procedure for the string theory computation which led to the desired match between gauge and string theory was presented.

The $AdS_4 \times CP^3$ set-up [12] allows one to consider a similar correlator i.e. an extremal three-point function involving two 1/2 BPS giant gravitons in combination with one 1/2 BPS point-like graviton and the methods developed in [5] for the AdS_5/CFT_4 calculation can be generalized to this case as well [13]. One remaining subtle point is the choice of regularization procedure in the string theory computation. In the $AdS_4 \times CP^3$ correspondence three-point functions of 1/2 BPS operators are not protected and hence in this set-up we cannot expect a match between gauge and string theory results. In particular, this means that on one hand we cannot justify our choice of regulator by a match between the gauge and string theory results but on the other hand a computation of the correlator in the weakly coupled string theory will provide us with a non-trivial prediction about the behaviour of the correlator in the dual strongly coupled field theory. Below we will revisit the regularization procedure employed for the $AdS_5 \times S^5$

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computation and modify it in a way that allows us to generalize it to the $AdS_4 \times CP^3$ case. Subsequently, we carry out the string theory calculation of the extremal three-point function involving two giant 1/2 BPS gravitons and one point-like 1/2 BPS graviton in $AdS_4 \times CP^3$.

1. Giant graviton correlators in $AdS_5 \times S^5$ revisited

Giant gravitons in $AdS_5 \times S^5$ are D3-branes which wrap an S^3 which constitutes a subset of either AdS_5 or S^5 [14–16]. We will consider the giants for which the wrapped sphere $S^3 \subset S^5$ and which spin along a circle in the S^5 while being located at the center of AdS_5 . For such giants the dual gauge theory operators are Schur polynomials built on completely anti-symmetric Young diagrams and containing a single complex scalar field that we will denote as Z [17,6]. The D3-brane action is (in units where the AdS radius has been set to 1)

$$S_{D3} = -\frac{N}{2\pi^2} \int d^4\sigma (\sqrt{-g} - P[C_4]), \tag{1}$$

where $g_{ab} = \partial_a X^M \partial_b X^N G_{MN}$, with $a, b = 0, \dots, 3$ being the world-volume coordinates and X^M the embedding coordinates. Working in global AdS coordinates

$$ds^2 = -\cosh^2 \rho dt^2 + d\rho^2 + \sinh^2 \rho d\tilde{\Omega}_3^2 + d\theta^2 + \sin^2 \theta d\phi^2 + \cos^2 \theta d\Omega_2^2, \tag{2}$$

the four-form potential C_4 can be written as [15]

$$C_{\phi\chi_1\chi_2\chi_3} = \cos^4 \theta \text{Vol}(\Omega_3), \tag{3}$$

where the χ_i are the angles of the wrapped sphere, i.e. $d\Omega_3^2 = d\chi_1^2 + \sin^2 \chi_1 d\chi_2^2 + \cos^2 \chi_1 d\chi_3^2$. Using the ansatz

$$\rho = 0, \quad \sigma^0 = t, \quad \phi = \phi(t), \quad \sigma^i = \chi_i, \tag{4}$$

one can show that a D3-brane with angular momentum k is stable when it sits at

$$\cos^2 \theta = k/N, \tag{5}$$

and spins at the speed of light, $\dot{\phi} = 1$ [14]. In order to determine the three-point function of two giant gravitons and a point-like graviton (i.e. a chiral primary) we should determine the variation of the Euclidean version of the D3-brane action in response to the insertion of the desired chiral primary at the boundary of AdS and subsequently evaluate these fluctuations on the Wick rotated giant graviton solution. As this procedure was described in detail in [5] we shall be brief here. Denoting the spherical harmonic representing the point-like graviton as Y_Δ (with Δ referring to its conformal dimension) we can write the variation of the D3-brane action as [5]

$$\delta S = \frac{N}{2\pi^2} \cos^2 \theta \int d^4\sigma \left(\frac{2}{\Delta+1} Y_\Delta (\partial_t^2 - \Delta^2) s^\Delta + 4 \left[\Delta \cos^2 \theta - \sin \theta \cos \theta \partial_\theta \right] s^\Delta Y_\Delta \right), \tag{6}$$

where s^Δ is the bulk to boundary propagator. As our spherical harmonic we choose

$$Y_\Delta = \frac{\sin^\Delta \theta e^{i\Delta\phi}}{2^{\Delta/2}}, \tag{7}$$

which corresponds to the single trace operator $\text{Tr} Z^\Delta$ in the gauge theory language. With this choice for Y_Δ the first term in eq. (6) is

finite and gives the following contribution to the three-point function structure constant

$$C_{finite}^3 = -\sqrt{\Delta} \frac{k}{N} \left(1 - \frac{k}{N} \right)^{\Delta/2}, \tag{8}$$

whereas the contribution coming from the term with square brackets takes the form of a divergent integral with a pre-factor which tends to zero. In Ref. [11] it was proposed to regularize the divergent integral by replacing the simple spherical harmonic Y_Δ with the more involved one

$$Y_{\Delta+2l,\Delta} = \mathcal{N}_{\Delta,l} \sin^\Delta \theta e^{i\Delta\phi} {}_2F_1(-l, \Delta+l+2; \Delta+1; \sin^2 \theta), \tag{9}$$

where $\mathcal{N}_{\Delta,l}$ is a normalization factor, and to consider the limit $l \rightarrow 0$ where $Y_{\Delta+2l,\Delta} \rightarrow Y_\Delta$, and where the contribution from the ill defined term of eq. (6) becomes finite. This procedure leads to a match between the string and gauge theory computation. The obtained match justifies the choice of the regularization procedure but does not suggest a general principle that one could build on when aiming at a generalization to giant gravitons in $AdS_4 \times CP^3$. One property which characterizes the spherical harmonic (9) is that it extends the simple one without making use of additional coordinates on S^5 . However, this property is somewhat deceptive and is not the correct clue to an extension to the $AdS_4 \times CP^3$ set-up.

Here we shall formulate the regularization procedure in a slightly different manner which will allow us to generalize it to the latter set-up. For that purpose we make use of the fact that spherical harmonics on S^5 are in one-to-one correspondence with symmetric traceless $SO(6)$ tensors. In particular (leaving out normalization factors) the spherical harmonic (7), which translates into $\text{Tr} Z^\Delta$ in the field theory, corresponds to the tensor¹

$$C_{\underbrace{1\dots 1}_{\Delta-k} \underbrace{2\dots 2}_k} = i^k, \tag{10}$$

where symmetrization is understood. It is easy to show that adding more indices of type 1 and type 2 to the tensor (i.e. adding more fields of type Φ_1 and Φ_2 to the operator) does not regularize the divergent integral. However, one can regularize the integral by considering the following symmetric traceless tensor

$$C_{\underbrace{1\dots 1}_{\Delta-k} \underbrace{2\dots 2}_k \underbrace{3\dots 3}_{2l-n} \underbrace{4\dots 4}_n} = i^{k+n}, \tag{11}$$

where $n < 2l$ and subsequently taking the limit $l \rightarrow 0$. Obviously, the gauge theory operator resulting from this tensor involves the complex field $Y = \Phi_3 + i\Phi_4$ in addition to the complex field Z . It is easy to check that the spherical harmonic (9) corresponds to an operator involving all six scalar fields of $\mathcal{N} = 4$ SYM but it is not straightforward to express the corresponding C-tensor in a closed form. The tensor (11) translates into the following spherical harmonic

$$Y_{\Delta l} = \mathcal{N}_{\Delta l} \sin^\Delta(\theta) e^{i\Delta\phi} \cos^2(l\theta) \sin^2(l\chi_1) e^{2il\chi_2}, \tag{12}$$

where $\mathcal{N}_{\Delta l}$ is another normalization constant. Using this spherical harmonic instead of Y_Δ when evaluating the second line of (6) and subsequently taking the limit $l \rightarrow 0$ gives us the following result for the regularized contribution to the three-point function

¹ The chiral primary operators of $\mathcal{N} = 4$ SYM can be written in the form $C_l^{i_1 i_2 \dots i_\Delta} \text{Tr}(\Phi_{i_1} \Phi_{i_2} \dots \Phi_{i_\Delta})$ where the Φ_i 's can be any of the six real scalar fields and where C_l is a symmetric traceless tensor. We take the complex scalar field Z to be given by $Z = \Phi_1 + i\Phi_2$.

$$C_{\text{regularized}}^3 = \frac{1}{\sqrt{\Delta}} \left(1 + \frac{k}{N} \Delta\right) \left(1 - \frac{k}{N}\right)^{\Delta/2}, \quad (13)$$

meaning that the total three-point function takes the form

$$C_{\text{total}}^3 = \frac{1}{\sqrt{\Delta}} \left(1 - \frac{k}{N}\right)^{\Delta/2}, \quad (14)$$

which precisely coincides with the gauge theory result obtained in [5]. One can also extend the C -tensor in eq. (11) to include indices of type 5 and 6 corresponding to the scalar fields Φ_5 and Φ_6 . This leads to the same result for the regularized contribution to the three-point function.

2. Giant graviton correlators in $AdS_4 \times CP^3$

The giant graviton that we will consider is most easily described in the M-theoretic language where it corresponds to an M5 brane which wraps two S^5 's intersecting at an S^3 , all inside S^7/\mathbb{Z}_m . Furthermore, the M5-brane rotates along a circle orthogonal to the compactification circle.² The general version of this brane was constructed in [23] but maximal versions appeared already in [21,22]. We shall be considering the general case for which an improved parametrization was found in [13] and we shall follow the notation of the latter reference. The field theory dual to this giant graviton is a Schur polynomial which involves two of the complex bi-fundamental scalar fields of ABJM theory and is built on the completely anti-symmetric Young tableau. For details we refer to [13,23], see also [18–20]. The giant graviton is a 1/2 BPS object and we will be interested in calculating the three-point function involving two such giants and a point-like graviton, likewise 1/2 BPS. The point-like graviton translates in the field theory language to a single trace operator involving the same two complex bi-fundamental fields as the giant. Unlike what is the case in the $AdS_5 \times S^5$ set-up, three-point functions of 1/2 BPS operators are not protected within the $AdS_4 \times CP^3$ correspondence and a match between tree level gauge theory and semi-classical string theory results is not to be expected. Thus the more reason there is to be careful about the choice of regulator for the string theory computations.

To describe the giant graviton it is convenient to use the parametrization of S^7/\mathbb{Z}_m introduced in [13]

$$Z_1 = r e^{\rho+i\theta} e^{i(\phi+\frac{\chi}{2})}, \quad Z_2 = r e^{-\rho+i\theta} e^{i(\phi-\frac{\chi}{2})}, \quad (15)$$

$$Z_3 = e^{\rho_3+i(\phi+\theta_3)}, \quad Z_4 = r_4 e^{i\phi}, \quad (16)$$

where $r_4^2 = 1 - e^{2\rho_3} - 2r^2 \cosh 2\rho$ and where the ranges of the coordinates are as follows

$$0 \leq r \leq \frac{1}{\sqrt{2}}, \quad \rho_3 \leq \frac{1}{2} \log(1 - 2r^2 \cosh 2\rho), \quad (17)$$

$$-\rho_{\max} \leq \rho \leq \rho_{\max} \quad \text{with} \quad \cosh 2\rho_{\max} = \frac{1}{2r^2}, \quad (18)$$

$$0 \leq \theta, \theta_3, \chi, m\phi \leq 2\pi. \quad (19)$$

The Z_i are then in one-to-one correspondence with the four complex bi-fundamental fields of ABJM theory. The giant graviton is described by the equation

$$Z_1 \bar{Z}_2 = \alpha^2 e^{i\chi(t)}, \quad (20)$$

where t is the global AdS time and α is a constant. Referring to the parametrization introduced in (15) and (16), the world volume coordinates of the brane are $(t, \rho, \rho_3, \theta, \theta_3, \phi)$ where

$$-\rho_M \leq \rho \leq \rho_M, \quad \cosh(2\rho_M) = \frac{1}{2\alpha^2}, \quad (21)$$

and the other ranges are as before. The brane is maximal for $\alpha = 0$ and minimal for $\alpha = \frac{1}{\sqrt{2}}$. The low energy effective action of the M5 giant is a sum of a DBI and a WZ term which in the given parametrization can be written as follows [13]

$$S_{\text{DBI}} = -c \cdot \alpha^2 \int d^6\sigma e^{2\rho_3} \sqrt{(\cosh 2\rho - 2\alpha^2\omega^2)(\cosh 2\rho - 2\alpha^2)},$$

$$S_{\text{WZ}} = c \cdot \alpha^2 \int d^6\sigma e^{2\rho_3} (\cosh 2\rho - 2\alpha^2)\omega, \quad (22)$$

where $c = R_{S^7/\mathbb{Z}_m}^6 / ((2\pi)^5 l_p^6)$ and $\omega = \dot{\chi}(t)$. As shown numerically in [23] the Routhian of the system is minimized for $\omega = 1$. Hence, like its $AdS_5 \times S^5$ cousin the present giant spins with the speed of light. The corresponding angular momentum k is related to its size in the following way [13]

$$\frac{k}{N} = \sqrt{1 - 4\alpha^4} - 4\alpha^4 \log \frac{1 + \sqrt{1 - 4\alpha^4}}{2\alpha^2}. \quad (23)$$

To calculate the three-point function involving two giant gravitons and a point-like graviton (a chiral primary) we should again evaluate the variation of the Euclidean action in response to the insertion of the chiral primary at the boundary of AdS . Denoting as before the spherical harmonic representing the chiral primary as Y_Δ we can write the variation of the action is

$$\delta S_{\text{DBI}} = c \cdot \int d^6\sigma \frac{1}{2} \sqrt{g} \left(2\Delta + \frac{\cosh 2\rho}{\cosh 2\rho - 2\alpha^2\omega^2} \left[\frac{4}{\Delta + 2} \partial_t^2 - \frac{\Delta^2}{\Delta + 2} \right] \right) Y_\Delta(\Omega) s^\Delta(X), \quad (24)$$

$$\delta S_{\text{WZ}} = -c \cdot \int d^6\sigma \omega \sqrt{g_{S^7/\mathbb{Z}_m}} g_{S^7/\mathbb{Z}_m}^{r\beta} \partial_\beta Y_\Delta(\Omega) \Big|_{r=\alpha} s^\Delta(X). \quad (25)$$

As our spherical harmonic we would like to use

$$Y_\Delta = \left(r^2 e^{i\chi} \right)^{\Delta/2}, \quad (26)$$

which corresponds to the single trace operator $\text{Tr}(Z_1 \bar{Z}_2)^{\Delta/2}$ containing the same fields as the giant graviton. With this choice of spherical harmonic the contribution to the three-point function from the term containing square brackets is finite and yields

$$C_{\text{finite}}^3 = \frac{1}{N} \left(\frac{\lambda}{2\pi^2} \right)^{1/4} (2k) (2\alpha^2)^\Delta \sqrt{2\Delta + 1}, \quad (27)$$

where $\lambda = N/m$. The remaining part of the contribution from eqs. (24) and (25) is ill-defined. In the same manner as it was the case for the $AdS_5 \times S^5$ set-up the ill-defined part takes the form of a divergent integral times a vanishing pre-factor. This again calls for a regularization of the contribution. With an eye to the regularization of Ref. [11] one would be tempted to look for a regulator in the form of a solution of the Laplace equation which would reduce to the simple one (26) when some of its quantum numbers were sent to zero and which would involve as few as possible additional coordinates compared to (26). However, the Laplace equation when expressed in the coordinates (15)–(16) can only in a few special cases be solved by separation of variables and the corresponding spherical harmonics turn out not to regularize the three-point function.

² For a detailed explanation of the $AdS_4 \times CP^3$ set-up we refer to [12].

On the other hand, the method based on using appropriate tensors as the starting point for the regularization, works neatly in the $AdS_4 \times S^7/\mathbb{Z}_m$ case as well and, in addition, allows us to understand why certain spherical harmonics fail to regularize the three-point function. The spherical harmonic (26) corresponds to the tensor³

$$C_{\Delta/2}^{\overbrace{1 \dots 1}^{\Delta/2}}_{\underbrace{2 \dots 2}_{\Delta/2}} = 1. \tag{28}$$

As in the $AdS_5 \times S^5$ case one can regularize the ill defined contribution to the three-point function by dressing the tensor above with extra indices corresponding to fields not already present in the chiral primary (26) and sending the number of extra indices to zero at the end of the calculation. A choice for the extension of the tensor indices which mimics closely the one of the $AdS \times S^5$ is

$$C_{\Delta/2}^{\overbrace{1 \dots 1}^{\Delta/2} \overbrace{3 \dots 3}^{l-k} \overbrace{4 \dots 4}^k}_{\underbrace{2 \dots 2}_{\Delta/2} \underbrace{3 \dots 3}_{l-k} \underbrace{4 \dots 4}_k} = (-1)^k, \tag{29}$$

where symmetrization is understood. This tensor corresponds to the spherical harmonic

$$Y_{\Delta l} = \tilde{N}_{\Delta l} r^\Delta e^{i\chi\Delta/2} e^{2l\rho} {}_2F_1(-l, -l; 1; e^{2\rho_3} r_4^2). \tag{30}$$

Using this spherical harmonic instead of Y_Δ from eq. (26) when evaluating the divergent part of (24) and (25) and sending l to zero at the end of the calculation leaves one with the following contribution to the three-point function

$$C_{regularized}^3 = -\frac{\Delta + 2}{\Delta} C_{finite}^3, \tag{31}$$

so that the full three-point function in the $AdS_4 \times \mathbb{CP}^3$ case becomes⁴

$$C_{total}^3 = \frac{1}{N} \left(\frac{\lambda}{2\pi^2} \right)^{1/4} (2k) (2\alpha^2)^\Delta \left(\frac{2\sqrt{2\Delta+1}}{\Delta} \right), \tag{32}$$

A natural question to ask is what would happen if we chose an even more general tensor as a regulator such as the tensor

$$C_{\Delta/2}^{\overbrace{1 \dots 1}^{\Delta/2} \overbrace{1 \dots 1}^{k_1} \overbrace{2 \dots 2}^{k_2} \overbrace{3 \dots 3}^{k_3} \overbrace{4 \dots 4}^{k_4}}_{\underbrace{2 \dots 2}_{\Delta/2} \underbrace{1 \dots 1}_{k_1} \underbrace{2 \dots 2}_{k_2} \underbrace{3 \dots 3}_{k_3} \underbrace{4 \dots 4}_{k_4}}} = (-1)^{k_2+k_4}, \tag{33}$$

where $\sum_{i=1}^4 k_i = l$. For such tensors it is possible to show that the regularization procedure leads to negative powers of α in the three-point function unless $k_1 = k_2$. Negative powers of α in the three-point function imply a divergence of the three-point function in the maximal limit and the correlator is thus not fully regularized. Enforcing $k_1 = k_2$ it is not possible to construct a traceless symmetric tensor unless $k_1 = k_2 = 0$ which brings us back to the previous case (29).

³ The chiral primary operators of ABJM theory can be written in the form $(C_l)_{j_1 \dots j_{\Delta/2}}^{i_1 \dots i_{\Delta/2}} \text{Tr}(Z_{i_1} \bar{Z}^{j_1} \dots Z_{i_{\Delta/2}} \bar{Z}^{j_{\Delta/2}})$ where the Z_i 's can be any of the four complex fields of the theory and where C_l is symmetric in upper and (independently) in lower indices and the trace taken over any pair of one upper and one lower index vanishes.

⁴ We have left out a sign here since three-point functions are anyway only determined up to a phase factor.

The regularization procedure suggested in [11] for extremal three-point functions involving two giant gravitons and a chiral primary in $AdS_5 \times S^5$ could be justified by its implication of an expected match between three-point functions of 1/2 BPS objects in string and gauge theory. As already mentioned, in the $AdS_4 \times \mathbb{CP}^3$ set-up three-point functions of 1/2 BPS operators are not protected and thus one has to rely on the analogy with the $AdS_5 \times S^5$ case when choosing a regularization procedure. By considering the procedure suggested in [11] from a slightly different angle we have been able to construct a generalization for the $AdS_4 \times \mathbb{CP}^3$ case. The procedure consists of dressing the tensor describing the chiral primary involved in the three-point function with indices of a type not already present, performing the calculation of the correlator and sending the quantum numbers counting the extra indices to zero. We notice that the outcome of this procedure is very similar to that of the $AdS_5 \times S^5$ case in that the extra contribution to the three point function coming from the regularized integral takes the form of a pre-factor times the finite part of the three-point function.

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References

- [1] N. Beisert, C. Ahn, L.F. Alday, Z. Bajnok, J.M. Drummond, L. Freyhult, N. Gromov, R.A. Janik, et al., *Lett. Math. Phys.* 99 (2012) 3, arXiv:1012.3982 [hep-th].
- [2] B. Basso, S. Komatsu, P. Vieira, arXiv:1505.06745 [hep-th].
- [3] Z. Bajnok, R.A. Janik, *J. High Energy Phys.* 1504 (2015) 042, arXiv:1501.04533 [hep-th].
- [4] E. D’Hoker, D.Z. Freedman, S.D. Mathur, A. Matusis, L. Rastelli, in: M.A. Shifman (Ed.), *The Many Faces of the Superworld*, World Scientific, 2000, pp. 332–360, arXiv:hep-th/9908160.
- [5] A. Bissi, C. Kristjansen, D. Young, K. Zoubos, *J. High Energy Phys.* 1106 (2011) 085, arXiv:1103.4079 [hep-th].
- [6] S. Corley, A. Jevicki, S. Ramgoolam, *Adv. Theor. Math. Phys.* 5 (2002) 809, arXiv:hep-th/0111222.
- [7] K. Zarembo, *J. High Energy Phys.* 1009 (2010) 030, arXiv:1008.1059 [hep-th].
- [8] R.A. Janik, P. Surowka, A. Wereszczynski, *J. High Energy Phys.* 1005 (2010) 030, arXiv:1002.4613 [hep-th].
- [9] M.S. Costa, R. Monteiro, J.E. Santos, D. Zoakos, *J. High Energy Phys.* 1011 (2010) 141, arXiv:1008.1070 [hep-th].
- [10] P. Caputa, R.d.M. Koch, K. Zoubos, *J. High Energy Phys.* 1208 (2012) 143, arXiv:1204.4172 [hep-th].
- [11] H. Lin, *J. High Energy Phys.* 1212 (2012) 011, arXiv:1209.6624 [hep-th].
- [12] O. Aharony, O. Bergman, D.L. Jafferis, J. Maldacena, *J. High Energy Phys.* 0810 (2008) 091, arXiv:0806.1218 [hep-th].
- [13] S. Hirano, C. Kristjansen, D. Young, *J. High Energy Phys.* 1207 (2012) 006, arXiv:1205.1959 [hep-th].
- [14] J. McGreevy, L. Susskind, N. Toumbas, *J. High Energy Phys.* 0006 (2000) 008, arXiv:hep-th/0003075.
- [15] M.T. Grisaru, R.C. Myers, O. Tafjord, *J. High Energy Phys.* 0008 (2000) 040, arXiv:hep-th/0008015.
- [16] A. Hashimoto, S. Hirano, N. Itzhaki, *J. High Energy Phys.* 0008 (2000) 051, arXiv:hep-th/0008016.
- [17] V. Balasubramanian, M. Berkooz, A. Naqvi, M.J. Strassler, *J. High Energy Phys.* 0204 (2002) 034, arXiv:hep-th/0107119.
- [18] T.K. Dey, *J. High Energy Phys.* 1108 (2011) 066, arXiv:1105.0218 [hep-th].
- [19] S. Chakraborty, T.K. Dey, *J. High Energy Phys.* 1203 (2012) 062, arXiv:1112.6299 [hep-th].

- [20] R. de Mello Koch, B.A.E. Mohammed, J. Murugan, A. Prinsloo, J. High Energy Phys. 1205 (2012) 037, arXiv:1202.4925 [hep-th].
- [21] D. Berenstein, D. Trancanelli, Phys. Rev. D 78 (2008) 106009, arXiv:0808.2503 [hep-th].
- [22] J. Murugan, A. Prinsloo, J. High Energy Phys. 1105 (2011) 129, arXiv:1103.1163 [hep-th].
- [23] D. Giovannoni, J. Murugan, A. Prinsloo, J. High Energy Phys. 1112 (2011) 003, arXiv:1108.3084 [hep-th].